STATISTICAL SURROGATE MODELS FOR PREDICTION OF HIGH-CONSEQUENCE CLIMATE CHANGE

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In safety engineering, performance metrics are defined using probabilistic risk assessments focused on the low-probability, high-consequence tail of the distribution of possible events, as opposed to best estimates based on central tendencies. We frame the climate change problem and its associated risks in a similar manner. To properly explore the tails of the distribution requires extensive sampling, which is not possible with existing coupled atmospheric models due to the high computational cost of each simulation. We therefore propose the use of specialized statistical surrogate models (SSMs) for the purpose of exploring the probability law of various climate variables of interest. An SSM is different than a deterministic surrogate model in that it represents each climate variable of interest as a space/time random field. The SSM can be calibrated to available spatial and temporal data from existing climate databases, e.g., the program for climate model diagnosis and intercomparison (PCMDI), or to a collection of outputs from a general circulation model (GCM), e.g., the community Earth system model (CESM) and its predecessors. Because of its reduced size and complexity, the realization of a large number of independent model outputs from an SSM becomes computationally straightforward, so that quantifying the risk associated with low-probability, high-consequence climate events becomes feasible. A Bayesian framework is developed to provide quantitative measures of confidence, via Bayesian credible intervals, in the use of the proposed approach to assess these risks.

KEY WORDS: Bayesian analysis, climate model, Karhunen-Loéve expansion, non-Gaussian random field, risk analysis

1. INTRODUCTION AND MOTIVATION

The US Climate Change Research Program's (CCRP) long-term performance measure is to "deliver improved scientific data and models about the potential response of the Earth's climate and terrestrial biosphere to increased greenhouse gas levels for policy makers to determine safe levels of greenhouse gases in the atmosphere." One interpretation of this goal is that: (1) there is a well-defined threshold above which levels of greenhouse gases are "unsafe;" (2) this threshold is possible to determine; and (3) the Earth's climate has not yet crossed the threshold. The CCRP meaning of "safe" is presumed to be the level "that would prevent dangerous anthropogenic interference with the climate system" referenced by the 1992 United Nations Framework Convention on Climate Change (UNFCCC) [1]. Schneider and Lane [2] proposed metrics for dangerous climate change, which spanned the sustainability measures of water, energy, health, agriculture, and biodiversity, and included risks associated with extreme weather events and irreversible cascading chains of events beyond "tipping points." As a consequence, we are motivated by metrics associated with high-consequence climate changes, which we define as the changes that would be experienced if the true climate sensitivity is in the upper tail of its underlying probability distribution. Climate sensitivity is commonly defined as the change in the global mean surface temperature after the climate system has reached a new equilibrium in response to a doubling of the CO₂ concentration in the atmosphere.

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The large and growing body of literature on global climate change is mostly written from a scientific perspective that focuses on the most probable future. A scientific approach is the most appropriate method for gaining understanding of natural systems by applying physically sound theory, empirical observations, and validated models. Scientifically conservative estimates are the ones that deviate the least from prior expectations. Scientific conservatism, when applied to climate change, tends to downplay the degree of change, and virtually all the climate change literature uses the term "conservative" in the opposite sense from that of safety engineers. The common approach is to generate probability density functions (PDFs) that encapsulate the best estimate of the future, plus some bounds on its uncertainty. The lower bound on expected climate change is the scientifically conservative estimate.

The Intergovernmental Panel on Climate Change (IPCC) reports present climate forecasts as assessments of the most probable future, with the tendency to err on the side of scientific conservatism. For example, the Fourth Assessment Report (AR4) of the IPCC provides a graph of "warming by 2090–2099 relative to 1980–1999 for non-mitigation scenarios" in terms of "best estimate and likely ranges of warming." "Likely" is defined by the AR4 as an outcome that occurs with a probability of more than 66%. Thus, the ranges provided by the IPCC for various scenarios tend to be of the most interest to decision makers because they are the most probable. Unfortunately, they are often treated as accurate forecasts to be used as the basis for informing policy decisions, as opposed to using high-consequence forecasts as suggested by Palmer [3].

The development of methods to quantify the uncertainty in climate sensitivity (and other climate system response characteristics) is a topic of ongoing research. Some assessments result in the generation of PDFs rather than the simple "likelihood bounds" as provided by the IPCC. These studies consistently show that the high-end sensitivities have a significant probability. For example, Forest et al., [4], give a 5%-to-95% confidence interval of 1.4°C to 7.7°C climate sensitivity with a distribution that is strongly skewed with a sharp cutoff at the low end and a fat tail at the high end. The sharp low-end cutoff is expected, because the best understood feedback in terms of normalized uncertainty (e.g., standard deviation divided by best estimate) is water vapor, and is strongly positive. Other feedbacks (clouds, albedo, and lapse rate) are not as well understood and have high uncertainty by comparison; see, for example, [5, Section 8.6.2.3]. Andronova and Schlesinger [6] calculated a distribution with a 10% probability of climate sensitivity greater than 6.8°C. Skewed distributions with a high-sensitivity tail are characteristic of climate sensitivity PDFs, and Roe and Baker [7] argue that such skewness is an inevitable consequence of the nature of the climate system and the inherent uncertainty in the feedbacks. More recent research suggests that the upper bound of the confidence interval in climate sensitivity may also have been underestimated [8] due to compensating feedbacks assumed in previous studies. This interpretation further suggests the need for a safety engineering approach to characterize the mechanisms that can lead to high sensitivity.

Murphy et al. [9] used the ensemble method in a "perturbed physics" approach, which systematically varied 29 model parameters to determine a probability distribution function that has a 5% to 95% range of 2.4°C to 5.4°C, with a median of 3.5°C and a most probable value of 3.2°C. This sophisticated analysis makes use of more advanced models and is consistent with the transient effects of climate change and forcing. Processes that determine climate sensitivity are varied systematically and uncertainties are weighted according to an objective index, described by Murphy et al., as "a measure of reliability that can be used to weight different GCMs according to the estimated relative likelihood that they will correctly predict climate change in the real world," were "GCMs" are general circulation models. However, the perturbed physics approach is prohibitively expensive in many cases due to the cost of the coupled climate model. In practice, one must construct a cheaper surrogate of a quantity of interest—say surface temperature or precipitation—whose samples are used to approximate statistics. Constructing a surrogate model from scattered data poses challenges for many existing techniques. In what follows, we propose a translation random field surrogate for a quantity of interest tuned to a set of output data from a coupled climate model. With this surrogate, we are able to explore tail probabilities of the quantity of interest that are essential to formal study of the consequences of the changing climate system.

This paper illustrates a general procedure for calibrating a statistical surrogate model to available data from large computer models for Earth climate. Our use of the phrase "climate change" does not necessarily imply a time-dependent model. The data used to calibrate the statistical models in the examples that follow are taken from specific time periods of a climate model. We treat temporal samples as independent samples when calibrating the model. The statistical model is therefore valid for the period of time indicated by the data and is not explicitly time-dependent.

However, one common way to characterize the magnitude of climate change is to "spin up" an equilibrium "doubled-CO₂" model run, and to determine what the climate would be with twice the preindustrial level of CO₂ compared to what it would be otherwise. The spin-up time, in which the climate is changing with time, is not usually the part of the problem that is of interest. It is the stable, time-independent "equilibrium" climate that receives the most study.

The outline of our paper is as follows. The statistical surrogate model (SSM) is developed in Section 2, with emphasis on calibration of the SSM to available data, the use of the SSM to make predictions, and the development of Bayesian methods to assess model credibility. The general framework is then applied to two different climate variables of interest in Section 3: global mean surface temperature and precipitation rate. The main purpose of our paper is to illustrate how statistical surrogate models and, in particular, translation random fields, can be calibrated to available data from large computer models for Earth climate. This has not been done before to our knowledge. We feel strongly that due to the ever-increasing size and complexity of climate models, the decision-making process cannot be fully integrated in the near term with these models, and the use of surrogates like the one proposed herein becomes necessary.

2. STATISTICAL SURROGATE MODEL FOR CLIMATE VARIABLES

Let $A(\lambda, \phi, z)$ denote a climate variable of interest, e.g., temperature, precipitation rate, total cloud fraction, etc., where $0 \le \lambda < 2\pi, -\pi/2 \le \phi < \pi/2$, and $z \ge 0$ are spatial coordinates denoting longitude, latitude, and geodesic altitude, respectively. To simplify notation, we use $\mathbf{u} = (\lambda, \phi, z)^T \in D = [0, 2\pi) \times [-\pi/2, \pi/2) \times [0, \infty)$ to collectively represent all spatial coordinates and simply write $A(\mathbf{u}), \mathbf{u} \in D$.

Oftentimes there is parametric uncertainty in climate model parameters, which propagates to model outputs. In addition, the chaotic nature of the equations describing climate physics can lead to variability in response. To represent this variability and uncertainty in climate variables, we model $A(\mathbf{u})$ as a random field, that is, $A(\mathbf{u})$ is a random variable for every fixed $\mathbf{u} \in D$. Hence, $A = A(\mathbf{u}, \omega)$ also depends on the additional argument $\omega \in \Omega$, where Ω represents the appropriate sample space of A. It is common practice to omit the explicit depiction of the functional dependence of A on ω and we shall do so here. Further, our convention is to use a capital letter or symbol to denote any random quantity; lowercase letters and symbols are reserved for deterministic quantities.

Our objective is to develop a class of surrogate models for A that can be used to make defensible predictions about the probability of various climate change scenarios of interest. Special emphasis is placed on the tails of the probability distribution, the calibration of the surrogate model to available data from current and future runs of complex climate models, and to providing a quantitative measure of confidence in any model predictions. Throughout the discussion we assume A is a scalar quantity; the approach can be extended to represent vector-valued climate variables.

2.1 Model Definition

Let $G(\mathbf{u})$ be a real-valued Gaussian random field with zero mean, unit variance, and covariance function $c(\mathbf{u}, \mathbf{v}) = \mathrm{E}[G(\mathbf{u})\,G(\mathbf{v})]$, where $\mathrm{E}[X]$ denotes the expected value of random variable X. We model A by a monotonic transformation of G of the following form

$$A(\mathbf{u}) = F_A^{-1} \circ \Phi \left[G(\mathbf{u}) \right], \tag{1}$$

where $F_A = F_A[a|\theta(\mathbf{u})]$ is an arbitrary cumulative distribution function (CDF) that depends on parameter vector $\theta \in \mathbb{R}^d$, and Φ is the CDF of a N(0,1) random variable, that is, a Gaussian or normal random variable with zero mean and unit variance. We note that the value for θ generally depends on spatial location $\mathbf{u} \in D$. Herein, we assume F_A is absolutely continuous, that is, we assume there exists an integrable function f_A such that $f_A(a|\theta) = \mathrm{d}F_A(a|\theta)/\mathrm{d}a$ is a probability density function (PDF). It can be shown that the random field A has marginal, or first-order, CDF F_A and marginal, or first-order, PDF f_A . That is, given the following finite dimensional distributions of random field $A(\mathbf{u})$:

$$F_m(a_1,\ldots,a_m;\mathbf{u}_1,\ldots,\mathbf{u}_m) = \Pr[(\mathbf{u}_1) \le a_1 \cap \cdots \cap A(\mathbf{u}_m) \le a_m],$$

the marginal CDF is [10, p. 375]

$$F_1(a_1, \mathbf{u}_1) = F_m(a_1, +\infty, \dots, +\infty; \mathbf{u}_1, \dots, \mathbf{u}_m).$$

This class of model defined by Eq. (1) is a special type of non-Gaussian random field referred to as a translation random field (see [11, Section 3.1.1] and [12]); it is common to refer to G as the Gaussian image of A. We note that the class of models defined by Eq. (1) forms a subclass of all non-Gaussian random fields, and is defined by the second-moment properties and marginal (or first-order) distribution function. The restriction of this class is that these models cannot represent information from higher-order distributions functions (second-order, third-order, and so on). However, robust estimates of these distributions are very difficult to obtain in many instances. Often, in our experience, information on the random field to be modeled is limited to second-moment properties and marginal distribution.

By careful selection of the properties of G, as well as the functional form of F_A and its parameters θ , it is possible to calibrate A to match statistical estimates of the mean, covariance, and marginal probability distribution functions of a climate variable of interest. The translation random field model defined by Eq. (1) is a very flexible model and has proved effective for a variety of applications in stochastic mechanics (see, for example, [13, 14]) and the modeling of various environmental phenomena [15, 16].

2.2 Model Calibration

Let $z_k(\mathbf{u})$, k = 1, ..., n, denote the available data on A, that is, results from a sequence of climate model calculations. We use $\mathbf{z} = [z_1(\mathbf{u}), ..., z_n(\mathbf{u})]^T$ to represent the collection of available data to simplify notation. Calibration of the random field model defined by Eq. (1) to \mathbf{z} requires three steps:

- 1. Choose the functional form for F_A , the marginal CDF of A defined by Eq. (1).
- 2. Calibrate θ , the associated parameters of F_A .
- 3. Construct the covariance function $c(\mathbf{u}, \mathbf{v})$ of G, the Gaussian image of A.

The objective of step 1 is to select a marginal distribution function F_A that is sufficiently flexible to capture any desired behavior observed in the data z, and is consistent with the known physics. For example, if A models precipitation rate, the distribution function must have support on the positive real line with positive skewness; the lognormal distribution satisfies these constraints and is often used to model precipitation rate [17].

Calibration of parameters $\theta = (\theta_1, \dots, \theta_d)^T$ to be consistent with available data \mathbf{z} , that is, step 2 of the calibration procedure, can be completed by one of two methods: the method of moments, or the method of maximum likelihood [18, Chapter 6]. For the former, we choose $\theta = \hat{\theta}$ such that $f_A(a|\hat{\theta})$ satisfies

$$\int_{\mathbb{R}} a^q f_A(a|\hat{\theta}) da = \frac{1}{n} \sum_{k=1}^n (z_k)^q, \ q = 1, \dots, d.$$
 (2)

For the latter, we choose $\theta = \hat{\theta}$ that maximizes

$$l(\mathbf{z}|\mathbf{\theta}) = \prod_{k=1}^{n} f_A(z_k|\mathbf{\theta}),\tag{3}$$

where $l(\mathbf{z}|\boldsymbol{\theta})$, viewed as a function of $\boldsymbol{\theta}$, represents the likelihood that the data have PDF $f_A(a|\boldsymbol{\theta})$. We note that by either method, the calibrated value for $\boldsymbol{\theta}$ generally depends on $\mathbf{u} \in D$.

Upon completion of steps 1 and 2 in the calibration procedure, the available data on A can be mapped to available data on G, i.e., $g_k(\mathbf{u}) = \Phi^{-1} \circ F_A[z_k(\mathbf{u})|\hat{\boldsymbol{\theta}}], k = 1, \ldots, n$. The objective for step 3 of the calibration procedure is to construct covariance function $c(\mathbf{u}, \mathbf{v}) = E[G(\mathbf{u}) G(\mathbf{v})]$ of G, a Gaussian random field with zero mean and unit

variance, that is consistent with the data $\{g_k(\mathbf{u}), k = 1, \dots, n\}$. One approach is to utilize the Karhunen-Loéve (K-L) representation for a Gaussian random field, i.e.,

$$G(\mathbf{u}) = \sum_{k>1} \sqrt{\zeta_k} \, \psi_k(\mathbf{u}) \, W_k, \, \mathbf{u} \in D, \tag{4}$$

where $\{\zeta_k, \psi_k(\mathbf{u}), k \geq 1\}$ are the eigenvalues and eigenfunctions, respectively, of $c(\mathbf{u}, \mathbf{v}) = E[G(\mathbf{u}) G(\mathbf{v})]$, and satisfy the integral equation

$$\int_{D} c(\mathbf{u}, \mathbf{v}) \, \psi_{k}(\mathbf{v}) \, d\mathbf{v} = \zeta_{k} \, \psi_{k}(\mathbf{u}), \tag{5}$$

and $\{W_k, k \geq 1\}$ is a collection of independent and identically distributed (iid) N(0,1) random variables. In practice, the infinite sum defined by Eq. (4) is truncated at $r \geq 1$ terms; we refer to the truncated sum as the K-L approximation for G. Calibration of the K-L approximation model requires a method to choose: (i) r, the number of terms retained in the sum, and (ii) the associated eigenvalues and eigenvectors, $[\zeta_k, \psi_k(\mathbf{u})], k = 1, \ldots, r$. An efficient method to do this based on the the singular value decomposition of a matrix containing the available data $\{g_k(\mathbf{u}), k = 1, \ldots, n\}$ is presented in Appendix A.

2.3 Model Prediction and Credibility

Let $h[A(\mathbf{u})]$ be a particular property of A that is of interest, e.g., a functional of climate field data such as the extreme or average value. Our objective is to predict values for

$$p(\boldsymbol{\theta}, \mathbf{u}) = \Pr \left\{ h[A(\mathbf{u})] \in S_h \mid \boldsymbol{\theta}(\mathbf{u}) \right\} = \int_{\{a: h(a) \in S_h\}} h(a) f_A(a|\boldsymbol{\theta}(\mathbf{u})) \, \mathrm{d}a, \tag{6}$$

where S_h denotes an appropriate "safe set," that is, a set such that if the quantity of interest $h[A(\mathbf{u})]$ remains within S_h , there is no cause for alarm. We note that it is oftentimes more useful to report values for 1-p which correspond to the probability that the quantity of interest departs from the safe set. The concept of a "tipping point" described in climate literature [19] is an example of the boundary of a safe set for this study.

Exact solutions to the integral defined by Eq. (6) do not exist in general. However, because generating a large number of samples from the surrogate model is computationally inexpensive, we can make use of straightforward Monte Carlo simulation to estimate the value for the integral in Eq. (6) with high accuracy. Further, as indicated by Eq. (6), our prediction p depends on the parameter vector θ which, due to limited data, is also uncertain. We can handle this issue in one of two ways, as described in Sections 2.3.1 and 2.3.2.

2.3.1 Point Estimates

The calibrated model for A can be used directly to provide point estimates for p defined by Eq. (6). Let \hat{p} denote an estimate for p. This estimate can be achieved by the following steps:

- 1. Generate $m \gg 1$ samples of the calibrated K-L approximation for G using a truncated version of Eq. (4).
- 2. Translate each sample of G using Eq. (1) with θ replaced by $\hat{\theta}$ defined in Section 2.2.
- 3. Evaluate the quantity of interest for each sample of A, herein denoted by h_k , $k = 1, \ldots, m$.

The point estimate for p defined by Eq. (6) is then given by

$$\hat{p} = \frac{1}{m} \sum_{k=1}^{m} \mathbf{1} (h_k \in S_h), \tag{7}$$

where $\mathbf{1}(\cdot)$ denotes the indicator function, i.e., $\mathbf{1}(B) = 1$ if event B is true and is zero otherwise. \hat{p} defined by Eq. (7) is an unbiased estimator for p and improves with increasing m. Hence, the point estimate can be made very accurate because generating many samples of the surrogate model (large m) is inexpensive.

2.3.2 Bayesian Approach

The approach of the previous section provides only point estimates for p defined by Eq. (6). We have limited confidence in these estimates because we have limited data on A, and we would like to quantify our level of confidence in some manner. One approach is to instead treat $\theta \in \mathbb{R}^d$, the parameters of the marginal CDF of A, as a random vector and apply a Bayesian approach. Bayesian credible sets can then be used to quantify prediction confidence.

Let Θ be a random vector with d coordinates and support $D_{\Theta} \subset \mathbb{R}^d$ representing the uncertain parameters of F_A defined by Eq. (1). Given a prior PDF, denoted by $f_{\Theta}(\theta)$, that describes our knowledge of Θ prior to studying the available data, the posterior PDF for Θ given the data is [20, Section 2.1]

$$f_{\Theta|\mathbf{z}}(\boldsymbol{\theta}|\mathbf{z}) \propto l(\mathbf{z}|\boldsymbol{\theta}) f_{\Theta}(\boldsymbol{\theta}),$$
 (8)

where $\mathbf{z} = [z_1(\mathbf{u}), \dots, z_n(\mathbf{u})]^T$ denotes the available data on the climate variable of interest, l denotes the likelihood function defined by Eq. (3), and the ∞ symbol is used to denote that the left and right sides of Eq. (8) are equal to within a normalizing constant.

With $\boldsymbol{\theta}$ replaced by $\boldsymbol{\Theta}$, the conditional probability $p(\boldsymbol{\Theta})$ defined by Eq. (6) is a random variable taking values in [0,1]. It is possible to find the probability that $p(\boldsymbol{\Theta})$ belongs to any subset C of [0,1]; this probability is equal to $\int_{C_{\boldsymbol{\Theta}}} p(\boldsymbol{\theta}) f_{\boldsymbol{\Theta}|\mathbf{z}}(\boldsymbol{\theta}|\mathbf{z}) \, d\boldsymbol{\theta}$, where $C_{\boldsymbol{\Theta}} = \{\boldsymbol{\theta} \in D_{\boldsymbol{\Theta}} : p(\boldsymbol{\theta}) \in C\}$. In particular, the $100(1-\alpha)\%$ credible set for $p(\boldsymbol{\Theta})$ is the set C such that [21, Section 4.3.2]

$$1 - \alpha \le \int_{\{\boldsymbol{\theta} \in D_{\boldsymbol{\Theta}}: \ p(\boldsymbol{\theta}) \in C\}} f_{\boldsymbol{\Theta}|\mathbf{z}}(\boldsymbol{\theta}|\mathbf{z}) \, \mathrm{d}\boldsymbol{\theta}. \tag{9}$$

For the special case when random variable $p(\Theta)$ has unimodal density, the credible set can be expressed as a credible interval, i.e., C = [a, b]. Additional constraints are, in general, needed to make the set C unique. For example, we can choose C such that the interval where the probability of being below the interval is as likely as being above it, or we can choose C such that the mean value is the central point. The latter is the constraint we will employ in the examples that follow.

The implementation of the Bayesian approach described in this section is as follows:

- 1. Postulate $f_{\Theta}(\theta)$, the prior PDF for model parameters Θ ; noninformative priors [20, Section 2.3] may be used for the case when no information on $f_{\Theta}(\theta)$ is available.
- 2. Compute the likelihood function and posterior PDF for Θ using Eq. (8).
- 3. Draw $m_1 \gg 1$ independent samples of random vector $\boldsymbol{\Theta}$ from $f_{\boldsymbol{\Theta}|\mathbf{z}}(\boldsymbol{\theta}|\mathbf{z})$; Markov chain Monte Carlo methods can be used for this step [22, 23].
- 4. For each sample θ_k of Θ , $k = 1, ..., m_1$, draw m_2 independent samples from $f_A(a|\theta_k)$.
- 5. Utilize the procedure in the previous section to estimate

$$\hat{p}_k = \frac{1}{m_2} \sum_{j=1}^{m_2} \mathbf{1}(h_{k,j} \in S_h), \tag{10}$$

where $h_{k,j}$ denotes the jth random sample drawn from $f_A(a|\theta_k)$.

6. Estimate the endpoints of the $100(1-\alpha)\%$ credible interval for $p(\Theta)$, denoted by $\hat{a} \leq \hat{b}$, where \hat{a} and \hat{b} satisfy

$$\frac{1}{m_1} \sum_{k=1}^{m_1} \mathbf{1}(\hat{p}_k \le \hat{a}) = \frac{\alpha}{2}$$

$$\frac{1}{m_1} \sum_{k=1}^{m_1} \mathbf{1}(\hat{p}_k > \hat{b}) = 1 - \frac{\alpha}{2}.$$
(11)

We conclude this section by noting that we have ignored any uncertainty in the parametric description of the Gaussian image of A, that is, in the eigenvalues and eigenvectors of the K-L approximation for G defined by Eq. (4). In principle, it is possible to include this information in the Bayesian approach outlined above if necessary.

3. APPLICATION TO CCSM DATA

To illustrate the use of the statistical surrogate model for climate variables developed in Section 2, we consider two collections of output from the Community Climate System Model (CCSM) v3.0, a fully coupled global climate model sponsored by the National Center for Atmospheric Research (NCAR). The first collection, presented in Section 3.1, corresponds to average December surface temperature for years 1990–1999 based on a collection of eight different model runs. We calibrate the surrogate model to the available model data and make various point predictions that may be of interest. The eight different model runs correspond to expert-tuned climate scenarios from the CCSM3. These models were included in the Coupled Model Intercomparison Project 3 [24], and their output is stored on the PCMDI multimodel database [25]. As is common in multimodel ensembles, we treat each run as an independent sample from an unknown distribution of climate scenarios for analysis purposes. The data from Section 3.1 are the average December surface air temperature from the years 1990 through 1999 for each of these eight runs. By the ergodic assumption commonly employed in climate models, we treat each year as a statistically independent sample. Therefore, in total we have 80 samples of the average December temperature over the globe for the period of 1990 through 1999.

In Section 3.2, we study average monthly precipitation rate over a 54-year period assuming a "cyclic Y2K ocean model." The observations are collected by averaging the precipitation rates over a month for June, July, and August; any variance in the year-to-year results is due to the chaotic nature of the model. We apply the calibrated surrogate model to study the probability that the precipitation rate falls below certain thresholds, then utilize the Bayesian approach described in Section 2.3.2 to quantify our confidence in these predictions.

3.1 Surface Temperature

In this section, the available data consist of eight different model runs over the 10-year period 1990 to 1999, as described above. We get one realization of the average December surface temperature field for each year, resulting in a total of 80 observations. Let $z_k(\mathbf{u})$, $k=1,\ldots,n=80$ denote the available surface temperature data; Fig. 1 illustrates the first data set $z_1(\mathbf{u})$ as contours of average surface air temperature as a function of longitude, λ , and latitude ϕ . The spatial grid is defined at lines of longitude and latitude with 1.4 degree spacing. Coordinates λ and ϕ

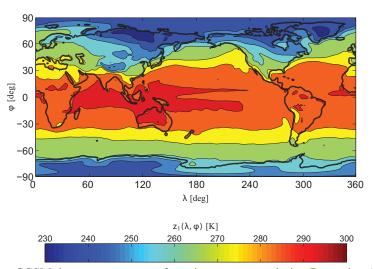


FIG. 1: First CCSM data set: average surface air temperature during December (from [25]).

representing longitude and latitude were discretized into 256 and 128 points, respectively, yielding a grid with 32,768 points. Upon inspection, we find that the data exhibit considerable skewness, i.e., the probability distribution of surface temperature is not symmetric about its mean value, and the degree of skewness changes with spatial location. An estimate of the coefficient of skewness is given by [26, p. 19]

$$\hat{\gamma}_3(\mathbf{u}) = \frac{n\sqrt{n-1}}{n-2} \frac{\sum_{k=1}^n \left[z_k(\mathbf{u}) - \hat{\mu}(\mathbf{u}) \right]^3}{\left(\sum_{k=1}^n \left[z_k(\mathbf{u}) - \hat{\mu}(\mathbf{u}) \right]^2 \right)^{3/2}}$$
(12)

where $\hat{\mu}(\mathbf{u}) = (1/n) \sum_{k=1}^{n} z_k(\mathbf{u})$ is an unbiased estimator for the mean surface temperature. The coefficient of skewness of the data is illustrated by Fig. 2. Regions of large positive skewness, e.g., parts of South America, correspond to locales where infrequent but large increases in temperature can be expected. It therefore seems appropriate to choose a non-Gaussian distribution that is able to match the large range of observed skewness exemplified by Fig. 2. In regions with near-zero skewness, a Gaussian distribution may be adequate.

Because of these observations, we choose F_A to be the CDF of a generalized version of the Gaussian distribution [27], that is, $F_A(a|\theta) = \int_{-\infty}^a f_A(\xi|\theta) d\xi$, where

$$f_A(a|\mathbf{\theta}) = \frac{1}{\sqrt{2\pi} \left[\theta_2 - \theta_3(a - \theta_1)\right]} e^{-y^2/2},$$
 (13)

is the corresponding PDF, $\theta = (\theta_1, \theta_2, \theta_3)^T$ is a vector of model parameters with support $\theta_1, \theta_3 \in \mathbb{R}$, $\theta_2 > 0$ and, assuming $\theta_3 \neq 0$,

$$y = -\frac{1}{\theta_3} \ln \left(1 - \frac{\theta_3(a - \theta_1)}{\theta_2} \right).$$

The CDF and PDF defined by Eq. (13) are illustrated by Figs. 3(a) and 3(b), respectively, for various values of parameters θ_1 , θ_2 , and θ_3 , which yields positive, negative, and zero skewness, denoted by $\gamma_3 > 0$, $\gamma_3 < 0$, and $\gamma_3 = 0$, respectively, in the figure. We apply the method of maximum likelihood to estimate values for the model parameters $\theta(\mathbf{u})$; see Section 2.2 and Appendix B.

For illustrative purposes, we make point predictions for two potential climate change scenarios of interest. First, we develop an estimate of

$$p_1 = \Pr\{h_1[A(\mathbf{u})] \in S_1\} = \Pr\left[\bar{A}(\mathbf{u}) < 2\right],\tag{14}$$

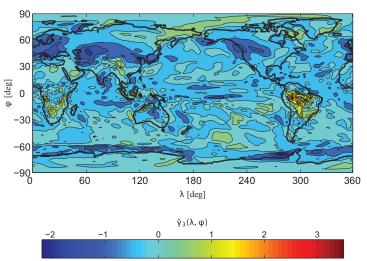


FIG. 2: Sample coefficient of skewness of all CCSM surface air temperature data.

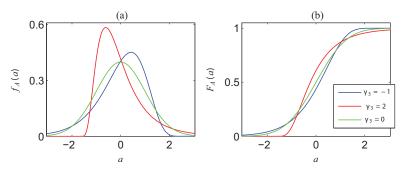


FIG. 3: Generalized normal distribution with positive ($\gamma_3 = 2$), negative ($\gamma_3 = -1$), and zero ($\gamma_3 = 0$) coefficients of skewness: (a) the PDFs, and (b) the CDFs. Each has zero mean and unit variance. For $\gamma_3 = 0$, F_A is a Gaussian distribution.

where $\bar{A}(\mathbf{u}) = \{A(\mathbf{u}) - \mathrm{E}[A(\mathbf{u})]\}/\sqrt{\mathrm{Var}[A(\mathbf{u})]}$ is a random field with zero mean and unit variance. By Eq. (14), $1 - p_1$ corresponds to the probability that A will exceed the so-called "2- σ level" and is illustrated by Fig. 4 as a function of latitude and longitude. This result illustrates a point estimate of the probability of higher than average air temperature during December, based on the available CCSM data and assumed model form. Since time has been omitted, these results should not be interpreted as a forecast of the likelihood of future warming.

Next let $D' \subset D$ be a region of interest whose boundary is defined by lines of longitude 170° and 240° and lines of latitude -30° and 30° (this corresponds roughly to the Pacific Ocean). Suppose the maximum average surface temperature within D' is a quantity of interest so that

$$p_2 = \Pr\left\{h_2[A(\mathbf{u})] \in S_2, \ \forall \mathbf{u} \in D'\right\} = \Pr\left(\max_{\mathbf{u} \in D'} A(\mathbf{u}) \le t\right). \tag{15}$$

Figure 5 illustrates a point estimate of $1-p_2$ as a function of temperature t; the estimate is based on 20,000 independent samples of A, the statistical surrogate model calibrated for average surface temperature.

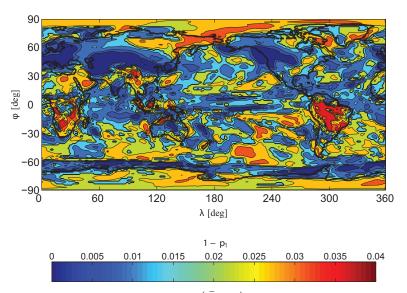


FIG. 4: Estimate of exceedance probability $1 - p_1 = \Pr(\bar{A} > 2)$ using the statistical surrogate model calibrated for data on air surface temperature.

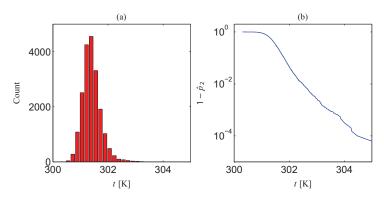


FIG. 5: Estimate of exceedance probability defined by Eq. (15) using the statistical surrogate model calibrated for air surface temperature: (a) histogram of 20,000 Monte Carlo samples of $\max_{\mathbf{u} \in D'} A(\mathbf{u})$, and (b) corresponding estimate for $1 - p_2$.

3.2 Precipitation Rate

The proposed statistical surrogate model has also been calibrated to CCSM model predictions of precipitation rate. This data set consists of monthly average precipitation rates for June, July, and August, over a 54-year period. Let $z_k(\mathbf{u}), k = 1, \ldots, n = 162$ denote the available data on precipitation rate; Fig. 6 illustrates $z_1(\mathbf{u})$ in units of mm/h as a function of longitude, λ , and latitude ϕ . As mentioned, the lognormal distribution is often used to model precipitation rate [17]; the corresponding PDF is given by

$$f_A(a|\theta_1, \theta_2) = \frac{1}{\sqrt{2\pi} \theta_2 a} e^{-(\ln a - \theta_1)^2/(2\theta_2^2)}, \ a > 0,$$
 (16)

where $\theta_1 \in \mathbb{R}$ and $\theta_2 > 0$ are model parameters.

We first consider point estimates of

$$p_3 = \Pr\left[\bar{A}(\mathbf{u}) \le -1\right],\tag{17}$$

where \bar{A} is defined by Eq. (14); p_3 corresponds to the probability that the precipitation rate will be equal to or less than the "1- σ level." Figure 7 illustrates an estimate of p_3 using the surrogate model calibrated to the available data;

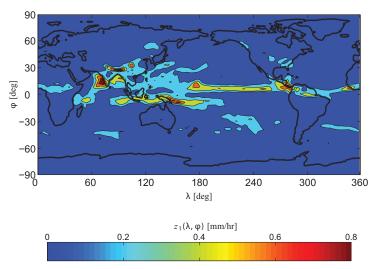


FIG. 6: First CCSM data set: precipitation rate during June, July, August.

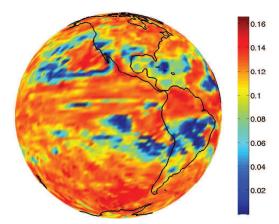


FIG. 7: Estimate of exceedance probability $p_3 = \Pr(\bar{A} \le -1)$ using the statistical surrogate model calibrated for data on precipitation rate.

this result illustrates a point estimate of the probability of drier than average conditions. Estimates of the model parameters $\hat{\theta}_1 = n^{-1} \sum_{k=1}^n \ln z_k$ and $\hat{\theta}_2^2 = n^{-1} \sum_{k=1}^n (\ln z_k - \hat{\theta}_1)^2$ are obtained using the method of maximum likelihood, Eq. (3).

A more useful result for the assessment of risk is to apply the Bayesian analysis described in Section 2.3.2. We predict

$$p_4(\rho) = \Pr\left[A(\mathbf{u}) \le \rho \right],\tag{18}$$

the probability that the precipitation rate is less than or equal to threshold ρ , and provide a measure of confidence on our predictions of p_4 as a function of ρ . To simplify the discussion, we present such an analysis for a fixed spatial location, i.e., a fixed $\mathbf{u} \in D$, that is roughly located at the city of Albuquerque, New Mexico. The analysis can be generalized to consider the entire domain D, but we focus here on a single grid point to clarify the discussion.

Following the procedure outlined in Section 2.3.2, we first model $\theta = (\theta_1, \theta_2)^T$, the parameters of the lognormal PDF defined by Eq. (16), as $\Theta = (\Theta_1, \Theta_2)^T$ a random vector with independent coordinates and noninformative prior PDF

$$f_{\Theta}(\theta_1, \theta_2) \propto \text{const}, \ \theta_1 \in \mathbb{R}, \theta_2 > 0.$$
 (19)

Given the available CCSM data on precipitation rate, **z**, application of Eqs. (8) and (16) yields the posterior PDF for the model parameters, i.e.,

$$f_{\mathbf{\Theta}|\mathbf{z}}(\theta_1, \theta_2|\mathbf{z}) \propto \theta_2^{-n} \exp\left(-\frac{n}{2\theta_2^2} \left(\hat{s}^2 - 2\hat{\mu}\theta_1 + \theta_1^2\right)\right), \ \theta_1 \in \mathbb{R}, \theta_2 > 0,$$
 (20)

where $\hat{\mu} = n^{-1} \sum_{k=1}^{n} \ln z_k$ and $\hat{s}^2 = n^{-1} \sum_{k=1}^{n} (\ln z_k)^2$ denote the sample mean and sample mean square of the natural log of the data. The posterior PDF for Θ is illustrated by Fig. 8(a).

Following step 3, the Markov chain Monte Carlo (MCMC) method [23] was used to draw 10,000 independent samples of random vector $\boldsymbol{\Theta}$ from $f_{\boldsymbol{\Theta}|\mathbf{z}}(\theta_1,\theta_2|\mathbf{z})$ defined by Eq. (20). These samples are illustrated by Fig. 8(b); level contours of the posterior PDF are also plotted to illustrate the performance of the MCMC method.

We next apply steps 4–6 from Section 2.3.2 to calculate the $100(1 - \alpha)$ % Bayesian credible interval for p_4 defined by Eq. (18); for calculations, we assume $\alpha = 0.1$. Two cases are studied, as illustrated by Figs. 9(a) and 9(b), respectively. In each case, the estimate of $p_4(\rho)$ is denoted by a blue solid line, while the corresponding 90% Bayesian credible intervals are illustrated by red dashed lines. In panel (a), we restrict the analysis to roughly half of the available data points. The resulting credible intervals are wider, indicating a lower degree of model credibility

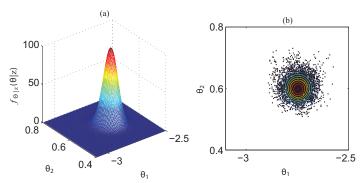


FIG. 8: Bayesian posterior analysis for precipitation rate data: (a) posterior PDF $f_{\Theta|\mathbf{z}}(\theta|\mathbf{z})$, and (b) 10,000 independent samples of Θ using MCMC. Level contours of $f_{\Theta|\mathbf{z}}(\theta|\mathbf{z})$ are also shown in panel (b).

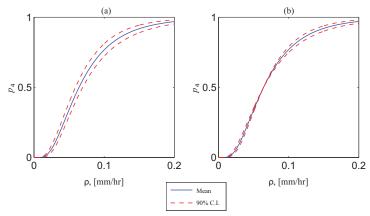


FIG. 9: Estimates of p_4 with 90% Bayesian credible intervals for the case of: (a) n = 78 data points, and (b) n = 162 data points.

when compared to panel (b), which corresponds to the same analysis based on the full data set. In addition, we note that the width of the credible intervals changes with ρ . Figures 7 and 9 illustrate the chance of drier than average conditions during June, July, and August, based on the predictions from the available data set and the assumed model form. Since time is not included, these results should not be interpreted as a forecast of the likelihood of future dry conditions.

As mentioned above, we used the method of maximum likelihood for model calibration. However, this approach provides a global fit to the data, rather than a localized fit to the data on extreme events, which may sometimes lead to inaccuracies in the tail regions of the SSM. One way to examine the accuracy of the construction in the tail region is to apply the method to synthetic data generated from a known density, then fit it with a different density, use the surrogate fitted density to estimate credible intervals, and compare to those of the true density. If accuracy is an issue, one could calibrate the model to a subset of the available data that better represents the tail regions.

4. CONCLUSIONS

We have formulated risk assessment in climate change impact studies in a framework similar to that used in safety engineering, by acknowledging that probabilistic risk assessments focused on low-probability, high-consequence climate events are perhaps more appropriate than studies focused simply on best estimates. To aid in this study, we have developed specialized statistical surrogate models (SSMs) that can be used to make predictions about the tails of the associated probability distributions. Without this approach, the use of models developed by the climate science

community for policy decisions will remain limited. The SSM represents each climate variable output of interest as a space/time random field, and can be calibrated to available spatial and temporal data from existing climate databases, or from a collection of outputs from global circulation models. The SSM used herein was a translation random field which, to the best of our knowledge, has not been used before to model Earth climate, and represents an advancement in applied uncertainty quantification (UQ) methods. Due to its reduced size and complexity, the realization of a large number of independent model outputs from an SSM becomes computationally straightforward, so that estimates of low-probability, high-consequence climate events becomes feasible. A Bayesian framework was also developed to provide quantitative measures of confidence, via Bayesian credible intervals, in the use of the proposed SSM as a statistical replacement for the associated GCM.

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APPENDIX A. CALIBRATION OF THE K-L REPRESENTATION FOR A GAUSSIAN RANDOM FIELD

We present a method to calibrate a covariance model of a Gaussian random field based on the singular value decomposition (SVD) of the data matrix of measurements described in Section 2.2. In practice, the data exist as point measurements over the spatial domain D; let $\mathbf{u}_j \in D$, $j=1,\ldots,m$ be the locations of measurements. We assume samples of the surrogate random field model occur at the \mathbf{u}_j . If we wish to sample the model at other points in D, we can use standard spatial interpolation techniques [28].

The kth measurement at the point \mathbf{u}_j is written $g_k(\mathbf{u}_j)$, which we assume comes from a zero mean random field. Define the $m \times n$ data matrix \mathbf{x} with entries

$$\mathbf{x}_{ik} = g_k(\mathbf{u}_i). \tag{A.1}$$

To calibrate an empirical covariance matrix, we compute the SVD of the zero-mean samples

$$\mathbf{x} = \mathbf{y} \,\mathbf{\sigma} \,\mathbf{v}^T. \tag{A.2}$$

We then pose the following surrogate model for the Gaussian random field $G(\mathbf{u}_i)$:

$$\begin{bmatrix} G(\mathbf{u}_1) \\ \vdots \\ G(\mathbf{u}_m) \end{bmatrix} \approx \frac{1}{\sqrt{n-1}} \sum_{i=1}^{n-1} \sigma_i \mathbf{y}_i W_i, \tag{A.3}$$

where \mathbf{y}_i are the left singular vectors and σ_i are the singular values. This is similar to the truncated Karhunen-Loéve expansion defined by Eq. (4), since \mathbf{y}_i are the eigenvectors of the empirical covariance matrix $[1/(n-1)]\mathbf{x}\mathbf{x}^T$, and $\sigma_i^2/(n-1)$ are its eigenvalues. Notice that the maximum possible truncation level is set by the number of available data n. However, we may be able to further truncate the expansion to less than n terms depending on the decay of the singular values. To generate samples of the calibrated field $G(\mathbf{u}_j)$, we draw samples from the iid Gaussian W_i and apply the transformation given by \mathbf{y} σ .

APPENDIX B. A GENERALIZED NORMAL DISTRIBUTION

Let A be a random variable with PDF $f_A(a|\theta_1, \theta_2, \theta_3)$ defined by Eq. (13). We can show that the support of the distribution, denoted by D_A , is given by

$$D_{A} = \begin{cases} (-\infty, \theta_{1} + \theta_{2}/\theta_{3}), & \theta_{3} > 0; \\ (\theta_{1} + \theta_{2}/\theta_{3}, \infty), & \theta_{3} < 0; \\ (-\infty, \infty), & \theta_{3} = 0, \end{cases}$$
(B.1)

meaning that the PDF defined by Eq. (13) is zero for all $a \notin D_A$. It follows that the mean, variance, and coefficient of skewness of A are

$$E[A] = \int_{D_A} a \, f_A(a) \, da = \theta_1 - \frac{\theta_2}{\theta_3} \left(e^{\theta_3^2/2} - 1 \right),$$

$$\operatorname{var}[A] = \int_{D_A} (a - E[A])^2 \, f_A(a) \, da = \left(\frac{\theta_2}{\theta_3} \right)^2 e^{\theta_3^2} \left(e^{\theta_3^2} - 1 \right),$$

$$\operatorname{skew}[A] = (\operatorname{var}[A])^{-3/2} \int_{D_A} (a - E[A])^3 \, f_A(a) \, da = \frac{3e^{\theta_3^2} - e^{3\theta_3^2} - 2}{\left(e^{\theta_3^2} - 1 \right)^{3/2}} \operatorname{sign}(\theta_3). \tag{B.2}$$

Further, A is a Gaussian random variable with zero skewness if, and only if, $\theta_3 = 0$; A is a non-Gaussian random variable with positive (negative) skewness when $\theta_3 < 0$ ($\theta_3 > 0$).

Given n independent samples of A, denoted by $\mathbf{z} = (z_1, \dots, z_n)^T$, the likelihood that this collection of samples was drawn from the generalized normal distribution is given by Eq. (3). This is often conveniently expressed as the negative of the log of the likelihood function, that is, $-\ln l(\mathbf{z}|\boldsymbol{\theta})$ where, by Eqs. (3) and (13)

$$-\ln l(\mathbf{z}|\boldsymbol{\theta}) = \frac{n}{2} \ln 2\pi + \sum_{k=1}^{n} \frac{1}{2\theta_3^2} \left[\ln \left(1 - \frac{\theta_3(z_k - \theta_1)}{\theta_2} \right) \right]^2 + \ln \left[\theta_2 - \theta_3(z_k - \theta_1) \right]. \tag{B.3}$$

It follows that the values for θ that maximize the likelihood function also minimize Eq. (B.3) and are the solution $\hat{\theta} = (\hat{\theta}_1, \hat{\theta}_2, \hat{\theta}_3)^T$ to the following set of equations:

$$\sum_{k=1}^{n} \frac{\hat{\theta}_{3}}{\hat{\theta}_{2} - \hat{\theta}_{3}(z_{k} - \hat{\theta}_{1})} + \frac{\ln \xi_{k}}{\hat{\theta}_{2} \, \hat{\theta}_{3} \, \xi_{k}} = 0$$

$$\sum_{k=1}^{n} \frac{1}{\hat{\theta}_{2} - \hat{\theta}_{3}(z_{k} - \hat{\theta}_{1})} + \frac{\ln \xi_{k}}{\hat{\theta}_{2}^{2} \, \hat{\theta}_{3} \, \xi_{k}} (z_{k} - \hat{\theta}_{1}) = 0$$

$$\sum_{k=1}^{n} \frac{\hat{\theta}_{1} - z_{k}}{\hat{\theta}_{2} - \hat{\theta}_{3}(z_{k} - \hat{\theta}_{1})} - \frac{(\ln \xi_{k})^{2}}{\hat{\theta}_{3}^{3}} - \frac{\ln \xi_{k}}{\hat{\theta}_{2} \, \hat{\theta}_{3}^{2} \, \xi_{k}} (z_{k} - \hat{\theta}_{1}) = 0,$$
(B.4)

where $\xi_k = 1 - (\hat{\theta}_3/\hat{\theta}_2)(z_k - \hat{\theta}_1)$.